### How does the near-horizon geometry of black rings encode its charges ?

Yuji Tachikawa (IAS)

#### in collaboration with K. Hanaki (U. Michigan) & K. Ohashi (DAMTP)

arXiv:0704.1819

Sep 21, 2007

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### 1. Introduction

2. Page charges

3. KK reduction

4. Summary

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• Open string d.o.f. accounts for Bekenstein-Hawking entropy for certain extremal Black Holes in 4D



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• Even subleading corrections from Wald's formula reproduced

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# Matching of the entropy



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# Matching of the entropy



#### The two methods give the same entropy !

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- Uniqueness theorem in 4D
- In 5D, the horizon can be  $S^3$  or  $S^2 imes S^1$ , Black Rings.





• More fun !

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- Entropy matching should be even more interesting.

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(a)

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- Entropy matching should be even more interesting.
- However ...



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• Hard to obtain Black Ring sol'n in asymptotically flat 5D

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- Harder with  $R^2$
- Hard enough in 4D. What made the analysis possible ?

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- Assumed BH is Extremal  $\longrightarrow$  Near-horizon is  $AdS_2 \times S^2$
- Electric Charge =  $\int_{S^2} \star F$ , Magnetic Charge =  $\int_{S^2} F$

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, Magnetic Charge =  $\int_{S^2} F$ 

• Assume BR is Extremal  $\longrightarrow$  Near-horizon is  $AdS_3 \times S^2$ 

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• Charges ??

### 5d minimal supergravity

$$S = rac{1}{8\pi G}\int \left(rac{1}{2}\star R - F\wedge \star F - rac{4}{3\sqrt{3}}A\wedge F\wedge F
ight)$$

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• 
$$d \star F = F \wedge F \longrightarrow \int_{\Sigma} \star F$$
 dependent on  $\Sigma$ 

- Electric charge distributes throughout the spacetime outside the near-horizon region
- Also true for angular momenta

### Page charge comes to the rescue ?

• 
$$d(\star F - A \wedge F) = 0$$
  
 $\longrightarrow \int_{\Sigma} (\star F - A \wedge F)$  independent on  $\Sigma$ 

• Called the Page charge

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• 
$$d(\star F - A \wedge F) = 0$$
  
 $\longrightarrow \int_{\Sigma} (\star F - A \wedge F)$  independent on  $\Sigma$ 

- Called the Page charge
- Not well-defined near horizon because
  - $\int_{S^2} F$  is non-zero
  - and thus A is not a globally well-defined one-form
- More complicated for angular momenta

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• We solve the problem by using two coordinate patches

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• We solve the problem by using two coordinate patches

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• BH, BR in  $\mathbb{R}^{3,1} \times S^1 \Longleftrightarrow$  BH in  $\mathbb{R}^{3,1}$ 

- We solve the problem by using two coordinate patches
- BH, BR in  $\mathbb{R}^{3,1} \times S^1 \iff$  BH in  $\mathbb{R}^{3,1}$
- Large gauge transformation  $\iff$  Witten effect on dyons

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### the Action

#### **Minimal Supergravity in 5D**

$$S = rac{1}{8\pi G}\int \left(rac{1}{2}\star R - F\wedge \star F - rac{4}{3\sqrt{3}}A\wedge F\wedge F
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#### EOM

$$egin{aligned} R_{\mu
u}&=-rac{1}{3}g_{\mu
u}F_{
ho\sigma}F^{
ho\sigma}+2F_{\mu
ho}F_{
u}{}^{
ho},\ d\star F&=-rac{2}{\sqrt{3}}F\wedge F. \end{aligned}$$

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# **Electric Charge**

$$F_{\mu\nu} = \begin{pmatrix} B_{ij} & E_i \\ -E_j & \mathbf{0} \end{pmatrix} \longrightarrow Q = \int_{\infty} d\vec{S} \cdot \vec{E} = \int_{\infty} \star F$$
  
defined at asymptotic infinity  
gauge-invariant and conserved  
sometimes called Maxwell charge

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$$\int_{\infty} \star F - \int_{\Sigma} \star F = \int_{B} d \star F = \int_{B} F \wedge F$$
$$\longrightarrow \rho^{\mu} = \epsilon^{\mu\nu\rho\sigma\tau} F_{\nu\rho} F_{\sigma\tau} \text{ is the electric current}$$

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charge diffusely distributed throughout the space

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# Page charge

• 
$$d \star F = F \wedge F = d(A \wedge F) \longrightarrow$$

$$\int_{\infty} (\star F - A \wedge F) = \int_{\Sigma} (\star F - A \wedge F)$$

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## Page charge

• 
$$d \star F = F \wedge F = d(A \wedge F) \longrightarrow$$

$$\int_\infty (\star F - A \wedge F) = \int_\Sigma (\star F - A \wedge F)$$

- introduced by D. Page
- $A \wedge F$  taken to vanish sufficiently fast near  $\infty \longrightarrow$

$$\int_{\infty} \star F = \int_{\infty} (\star F - A \wedge F)$$

• invariant under small gauge transformation  $A 
ightarrow A + d\chi$  because

$$\int_{\Sigma} (A + d\chi) \wedge F = \int_{\Sigma} A \wedge F + d(\chi \wedge F) = \int_{\Sigma} A \wedge F.$$

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• Not invariant under Large gauge transformation  $A \rightarrow A + g^{-1}dg$ where  $g = \exp(i\phi)$  winds nontrivially along some path C

$$\int_{C \times S^2} (g^{-1} dg) \wedge F = (\phi(2\pi) - \phi(0)) \int_{S^2} F$$
$$= (2\pi)^2 \times \text{integer}$$

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• can be fixed at asymptotic infinity:  $g^{-1}dg$  only decay  $\sim 1/r$ 

## Gauge Non-invariance of Page charge

• cannot really be fixed if there are multiple black objects :

$$\underbrace{\int_{\Sigma} (\star F - A \wedge F)}_{Q} = \underbrace{\int_{\Sigma_{1}} (\star F - A \wedge F)}_{Q_{1}} + \underbrace{\int_{\Sigma_{2}} (\star F - A \wedge F)}_{Q_{2}}$$

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• Q can be fixed, but  $Q_1$  and  $Q_2$  might change.

### Even worse.

• **A** not a globally well-defined one-form if 
$$\int_{S^2} \neq 0$$

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$$\rightarrow \int A \wedge F$$
 meaningless.

• Way out :

0

**1** Introduce the Dirac surface.

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: extends outside near-horizon region.

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#### **2** Introduce two coordinate patches.

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### **Solution**



Let  $A_S = A_T + \beta$  where  $\beta = g^{-1}dg$   $\longrightarrow$ 

$$\int_{B} F \wedge F = \int_{B \cap S} F \wedge F + \int_{B \cap T} F \wedge F$$
  
= 
$$\int_{C \cap S + D \cap B} A_{S} \wedge F + \int_{C \cap T - D \cap B} A_{T} \wedge F$$
  
= 
$$\left(\int_{C \cap S} A_{S} \wedge F + \int_{C \cap T} A_{T} \wedge F\right) + \int_{D \cap B} (A_{S} \wedge F - A_{T} \wedge F)$$
  
= 
$$\left(\int_{C \cap S} A_{S} \wedge F + \int_{C \cap T} A_{T} \wedge F\right) + \int_{C \cap D} A_{S} \wedge \beta.$$

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### **Solution**

• Define

$$^{\prime\prime}\int_{M}A\wedge F\stackrel{\prime\prime}{=}\int_{M\cap S}A_{S}\wedge F+\int_{M\cap T}A_{T}\wedge F+\int_{D\cap M}A_{S}\wedge \beta$$

$$\int_{\infty} (\star F - A \wedge F) = \int_{\Sigma} (\star F - A \wedge F)$$
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where you can take  $\boldsymbol{\Sigma}$  to be the horizon.
### **Solution**

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$${}^{\prime\prime}\int_MA\wedge F\,{}^{\prime\prime}\equiv\int_{M\cap S}A_S\wedge F+\int_{M\cap T}A_T\wedge F+\int_{D\cap M}A_S\wedge eta$$

• Then 
$$\int_{\infty} (\star F - A \wedge F) = \int_{\Sigma} (\star F - A \wedge F)$$

where you can take  $\boldsymbol{\Sigma}$  to be the horizon.

• Still non-invariant under large gauge transformation.

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- Assume spacetime to be stationary, let *ξ* be Killing vector generating rotation
- Komar's formula

$$J_{\xi} = -\frac{1}{16\pi G} \int_{\infty} \star \nabla \xi$$

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where  $\nabla \xi = \nabla_{\mu} \xi_{\nu} dx^{\mu} \wedge dx^{\nu} = d\xi.$ 

•  $d \star d\xi = 2 \star R_{\mu\nu}\xi^{\mu}dx^{\nu} \longrightarrow$ Komar's integral independent of the surface for vacuum gravity not independent otherwise.

### Wald's Noether charge

• Wald's analysis guarantees one can always find

$$d \star \nabla \xi = 2 \star R_{\mu\nu} \xi^{\mu} dx^{\nu}$$
  
=  $4 \star F_{\mu\rho} F^{\rho}{}_{\nu} \xi^{\mu} dx^{\nu} = d$ (something)

so that

$$\int_{\Sigma} (\star \nabla \xi + \text{something})$$

is independent of  $\Sigma$ .

• that "something" is

$$\star(\xi\cdot A)F+\frac{4}{3\sqrt{3}}(\xi\cdot A)A\wedge F.$$

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Not well-defined because
 A is not globally well defined at the horizon.

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- can be made well-defined using two patches as before.
- rather than being concrete, let us give the general prescription:

 $0 \ d\omega = \rho \text{ where } \rho = F \wedge F \text{ and/or } 2 \star R_{\mu\nu} \xi^{\mu} dx^{\nu}.$ 

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 $\bigcirc$  where, to compensate,  $d\omega^{(1)} = \omega_S - \omega_T$ 

**6** which can be solved because

$$dd\omega^{(1)} \stackrel{?}{=} d(\omega_S - \omega_T) = \rho_S - \rho_T = \mathbf{0}$$

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**6** which is the Descent equation !

Checks

Supersymmetric Single Black Ring:



$$egin{aligned} A_S &= -\left(q+rac{Q}{q}
ight)d\psi - q\cos heta d\chi, \ A_T &= A_S + 2qd\phi \end{aligned}$$

electric charge

$$= \int_{S \cap \Sigma} A_S \wedge F + \int_{D \cap \Sigma} A_T \wedge A_S$$
$$= \frac{Q + q^2}{2} + \frac{Q - q^2}{2} = Q$$

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• Supersymmetric Multiple Black Rings

$$\mathbf{Q}_{\text{total}} = \frac{\sqrt{3}\pi}{2G} \left[ \sum_{i=1}^{N} \left( Q_i - q_i^2 \right) + \left( \sum_{i=1}^{N} q_i \right)^2 \right],$$

where Q<sub>i</sub> and q<sub>i</sub> are the electric and magnetic charges of *i*-th ring.
"Interaction" comes from the compensating term

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- We considered BR in  $\mathbb{R}_t \times \mathbb{R}^4$
- Let us instead consider BR in  $\mathbb{R}_t \times$  Euclidean Taub-NUT



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Image: Image:

• Near the NUT, it's just a BR in  $\mathbb{R}^{4,1}$ 

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- Near infinity, it's just a Black String wound in KK circle

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- Near the NUT, it's just a BR in  $\mathbb{R}^{4,1}$
- Near infinity, it's just a Black String wound in KK circle
- Near-horizon geom. the same in proper coordinates

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## Gauge dependence



• 
$$a=\int A_\psi d\psi$$
 is a scalar field in 4d

- "a(x) 
  ightarrow 0 when  $|x| 
  ightarrow \infty$ " is not the right thing
- large gauge transformation  $A \to A + g^{-1}dg$  $\longrightarrow a \to a + 2\pi n$
- charge q, electric charge Q, angular momentum J changes

$$(q,Q,J) 
ightarrow (q,Q+q,J+q^2+qQ)$$

because 
$$Q \supset \int AF$$
,  $J \supset \int AAF$ 

## **KK reduction**

$$S_{5D} = \frac{1}{2} \star R - F^{I} \wedge \star F^{J} - \frac{4}{3\sqrt{3}}A \wedge F \wedge F^{K}$$

$$ds_{5d}^{2} = e^{2\rho}(d\psi - A^{0})^{2} + e^{-\rho}ds_{4d}^{2}$$

$$ds_{5d} = a(d\psi - A^{0}) + A_{4d}^{1}$$

$$S_{gauge fields} = \left[\frac{1}{2}e^{3\rho} + 2e^{\rho}a^{2}\right]F^{0} \wedge \star F^{0} - \left(\frac{2a}{\sqrt{3}}\right)^{3}F^{0} \wedge F^{0}$$

$$+ 4e^{\rho}aF^{0} \wedge \star F^{1} + \frac{4}{\sqrt{3}}a^{2}F^{0} \wedge F^{1}$$

$$- 2e^{\rho}F^{1} \wedge \star F^{1} - \frac{4}{\sqrt{3}}aF^{1} \wedge F^{1}$$

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- Just kinetic term +  $\theta$  term

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- Detailed structure doesn't matter much
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- Electric and magnetic charge well-defined,
- no problem associated to Chern-Simons
- What happened to the gauge-dependence ?
- $\epsilon^{\mu\nu\rho\sigma}\theta F_{\mu\nu}F_{\rho\sigma}$  with scalar dependent  $\theta$

$$-\left(rac{2a}{\sqrt{3}}
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#### Witten Effect

• Consider 
$$\mathcal{L} = F_{\mu\nu}F^{\mu\nu} + \theta\epsilon^{\mu\nu\rho\sigma}F_{\mu\nu}F_{\rho\sigma}$$

•  $\theta$  doesn't change the EOM, but changes

$$B_{i} = \frac{1}{2} \epsilon_{ijk} F^{jk},$$
  

$$E_{i} = \frac{\partial \mathcal{L}}{\partial B_{i}} = E_{i} + \theta B_{i}.$$

$$Q = Q_{0} + \theta q$$

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- i.e.  $\theta \rightarrow \theta + \epsilon$  changes the physics
- but still  $\theta \sim \theta + 2\pi$  equivalent

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- but still  $\theta \sim \theta + 2\pi$  equivalent
- Adiabatic change  $\theta \rightarrow \theta + 2\pi$  brings  $(Q,q) \rightarrow (Q + 2\pi q,q)$

#### **5d vs 4d**

• in 5d, large gauge tr.  $A \rightarrow A + g^{-1}dg$  changes

$$(q,Q,J) \longrightarrow (q,Q+q,J+Qq+q^2)$$

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 $a \rightarrow a + 2\pi$  changes

 $(q_1, Q^1, Q^0) \longrightarrow (q_1, Q^1 + q_1, Q^0 + q_1Q^1 + (q_1)^2)$ 

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•  $Q^0$ : KK electric charge = J: angular momentum along  $\psi$ •  $a = \int A_{\psi} d\psi$ 

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- String theoretic construction of Black Rings from M2-branes and M5-branes and momentum on them
- q: number of M5 branes

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$$ar{Q}$$
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• Equivalent under large gauge transformation / Witten effect.

## **Final example**



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## **Final example**



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## **Final example**



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1. Introduction

2. Page charges

3. KK reduction

## 4. Summary

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- Two notions of charge: Maxwell vs. Page
- Careful definition of Page charge: two coordinate patches with a boundary contribution

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- Two notions of charge: Maxwell vs. Page
- Careful definition of Page charge: two coordinate patches with a boundary contribution

Conserved charges encoded in the near-horizon geometry !

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Large gauge transformation in 5d
 ↔ Witten effect on dyons in 4d

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- Large gauge transformation in 5d
  ↔ Witten effect on dyons in 4d
- Higher-derivative interaction ?? [Castro-Davis-Kraus-Larsen]
- NUT gets electric charges from  $A \land R \land R$ ??

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